

SENR/NRPy+: Numerical Relativity in Singular, Curvilinear Coordinate Systems

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In collaboration with
Ian Ruchlin 

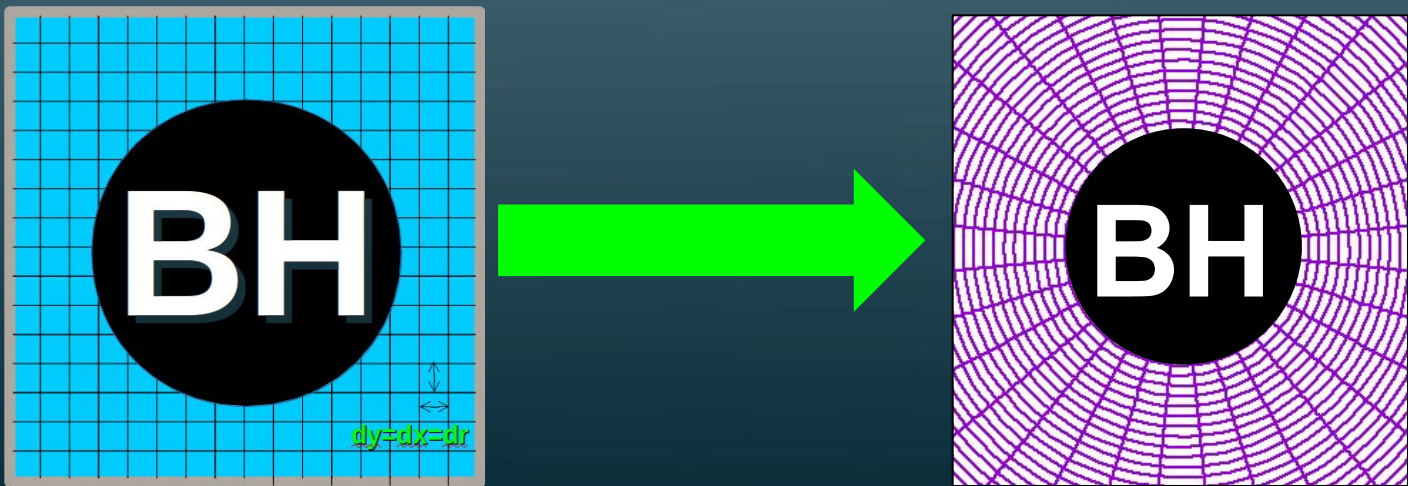
Thomas W. Baumgarte 

Phys. Rev. D 97, 064036 (2018)

<http://tinyurl.com/SENRcode>

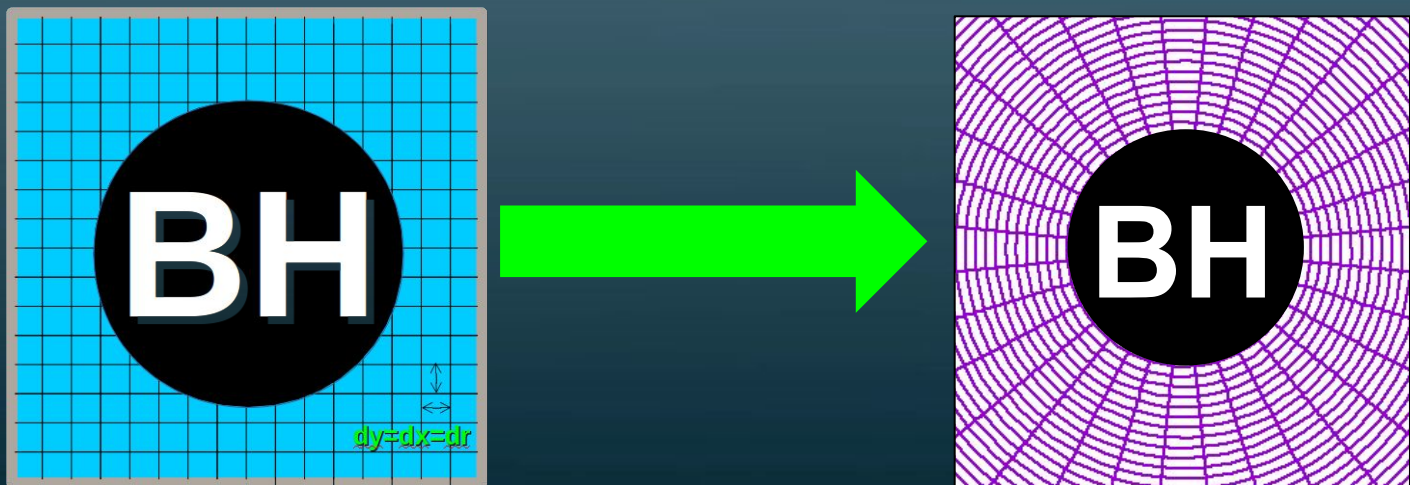
Beyond Black Holes in a Box: Exploiting Near-Symmetries in NR with Singular Coords

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Exploit near-symmetries → Minimize sampling points

BHB Idea: Spherical coord. sampling near each BH & far away

BoE: ~180x fewer points for BHBs than Cartesian AMR

Supercomputer → Consumer-grade desktop!

Toward Black Hole Binaries on the Desktop

1. Extend BSSN-in-spherical-coordinates technique to broader class of static, singular coordinates
 - Implement this technique in new **SENr/NRPy+** code



This talk

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2. Singular coordinates that comove with binary
 - First moving puncture BHB evolutions in singular curvilinear coords
 - First BHB evolutions in corotating frame, without excision



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Addressing Issues with Singular Coordinates

1. Tensor components can be **singular** ($\rightarrow 0$ or ∞) at coord singularities
 - Singular pieces are multiplicative and known analytically
 - Scale out singular pieces & handle spatial derivs analytically
 - Promote rescaled tensors to evolved quantities

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2. Divergent multiplicative terms in RHSs of equations

- E.g., 1D scalar wave equation:

$$\partial_t^2 u = \partial_r^2 u + \frac{2}{r} \partial_r u$$

- $2/r$ term “stiffens” the equation
- Even with cell-centered grids, RK2 timestepping is unstable
 - Can use PIRK2 (original formulation), but
 - **Ordinary RK4 works just fine in 3+1 NR** (discovered later)

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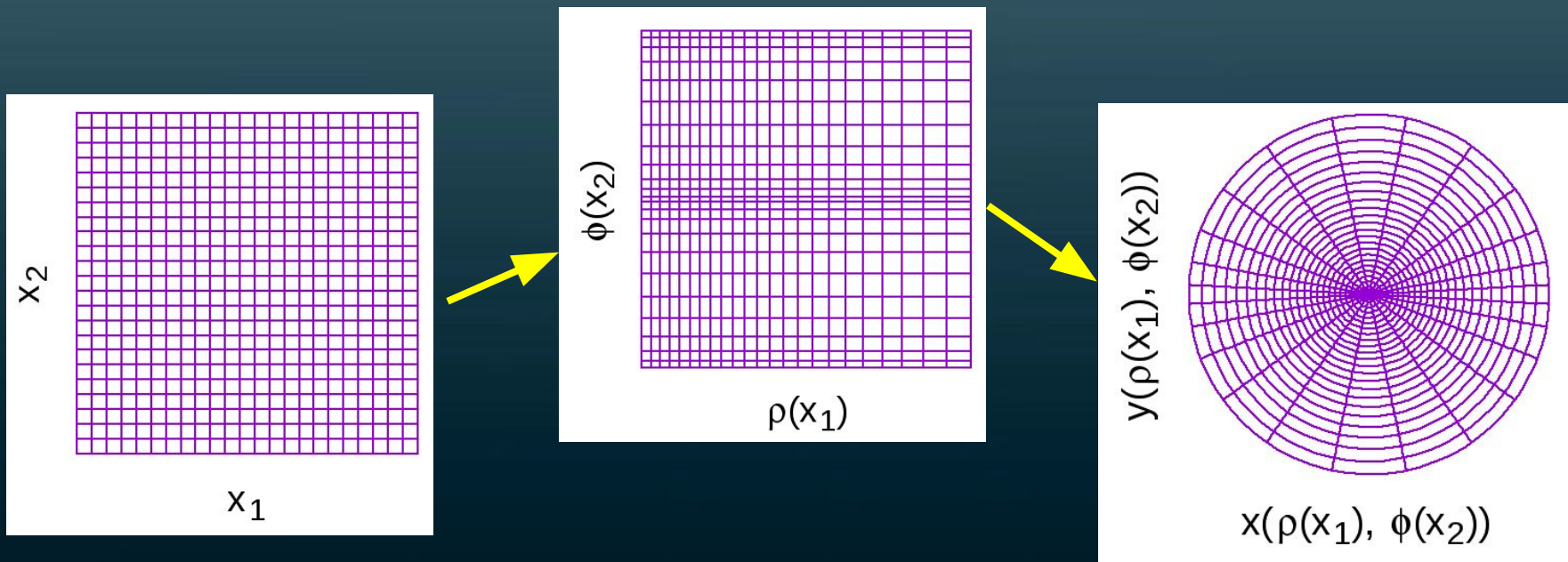
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**Net result: Stability & convergence properties
on par with Cartesian grids**

SENR/NRPy+: BSSN in Singular Coordinate Systems

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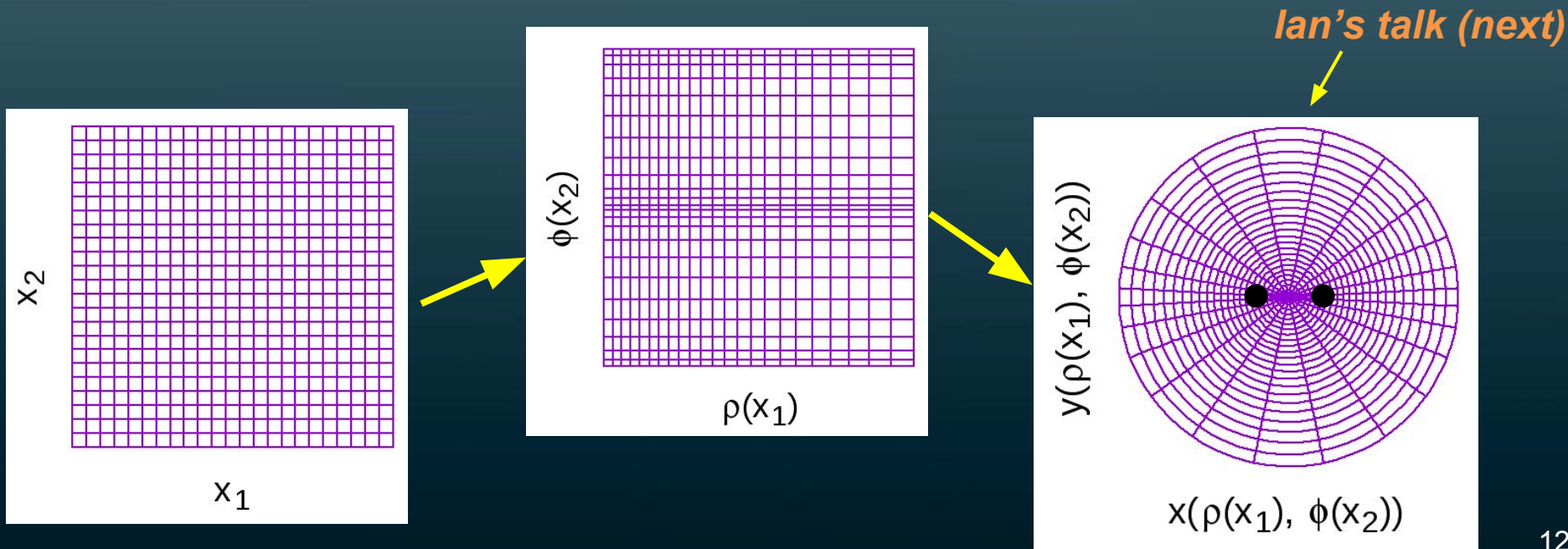
- Our work *Phys. Rev. D* 97, 064036 (2018), arXiv: 1712.07658
 - Generalize BSSN moving puncture formalism, to much broader class of singular coordinate systems



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SENR/NRPy+: Code Overview

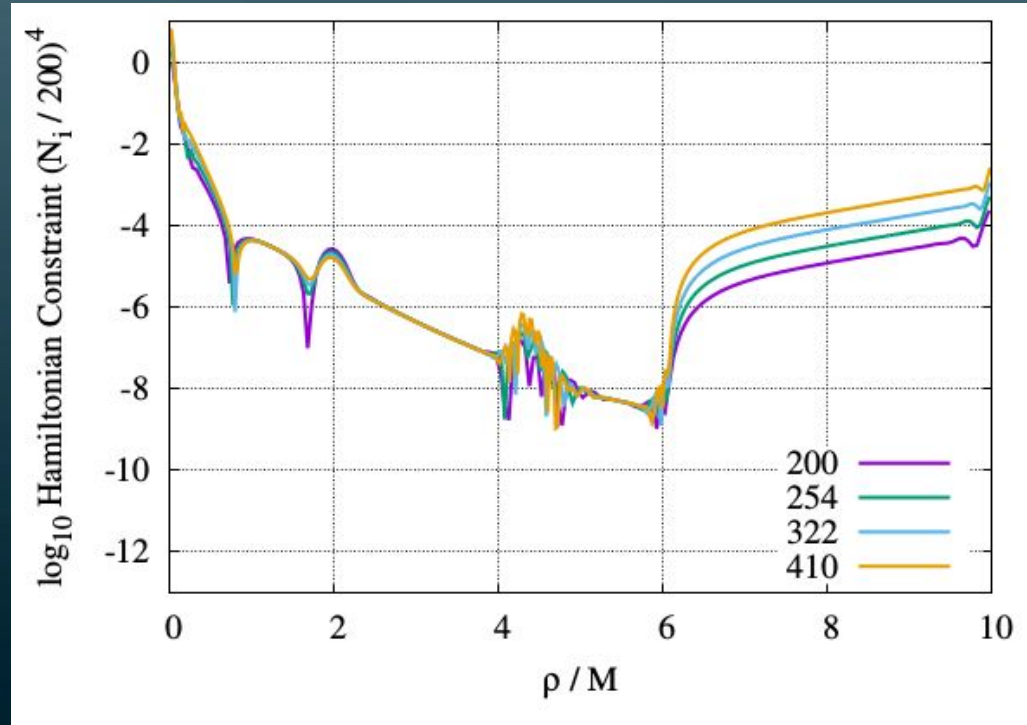
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- **NRPy+:** “Python-based code gen for NR... and beyond!”
 - Input: Python expressions in Einstein notation
 - Output: Highly efficient C code (AVX intrinsics + FD stencil capable)
 - Implements BSSN in spherical-like, cylindrical-like, and Cartesian-like coords
- **SENR:** “Simple, Efficient Numerical Relativity code”
 - Everything an OpenMP-parallelized NR code needs except equations
 - NRPy+ provides needed C code
 - SENR: RK4 timestepping, boundary condition routines, etc.
 - In spherical coords, results agree to roundoff with Baumgarte *et al* code
- **Both:** Open source (BSD-licensed), designed for user friendliness

SENR/NRPy+: Code Validation

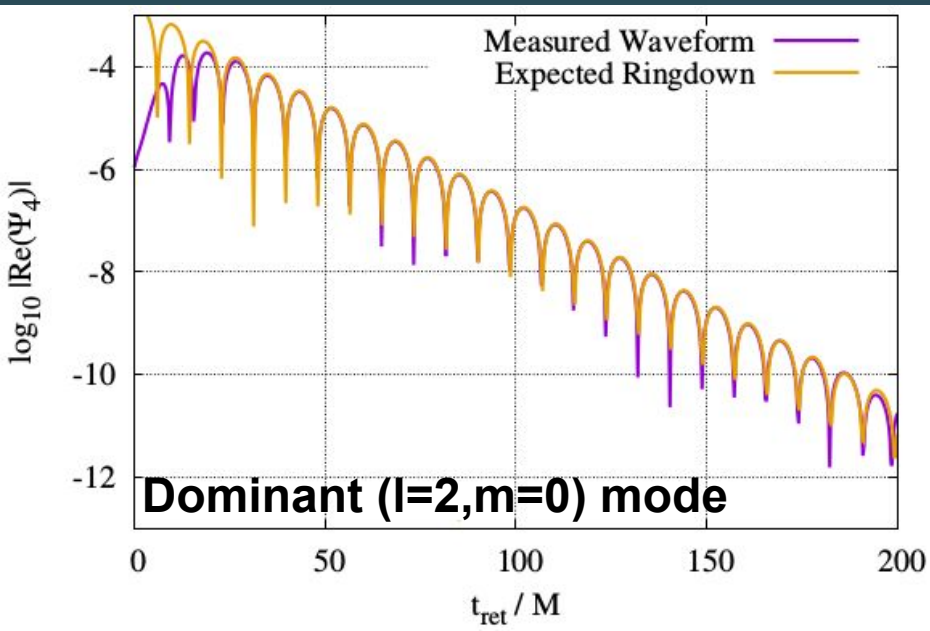
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- Black hole simulation
 - Wormhole initial data
 - Cylindrical coordinates
 - Fourth-order finite differencing
- Excellent convergence
 - at $t = 5M$, in region unaffected by outer boundary (at $r=10M$)



BH Spectroscopy from **Head-on BH Collision**

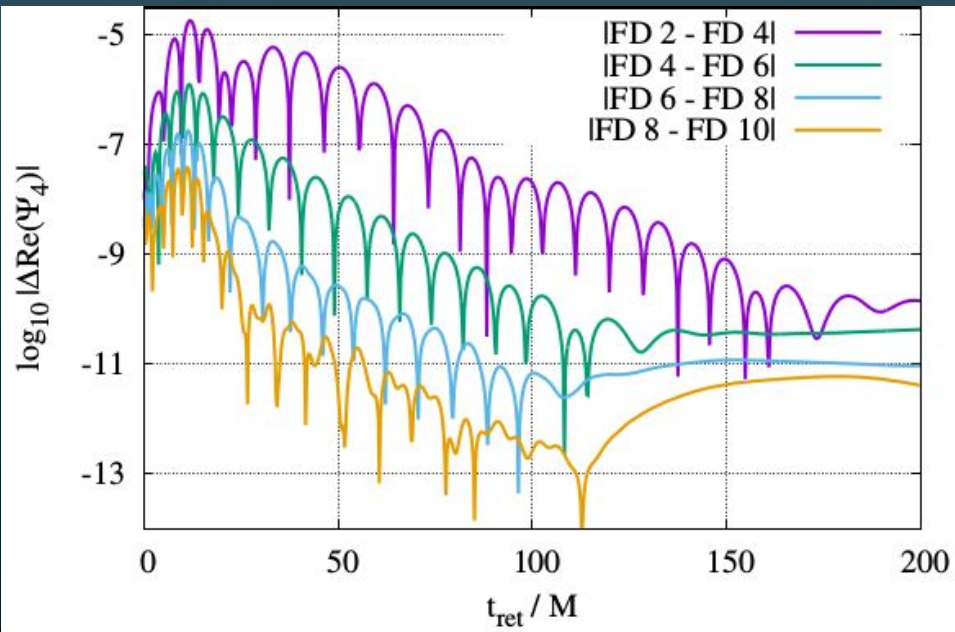
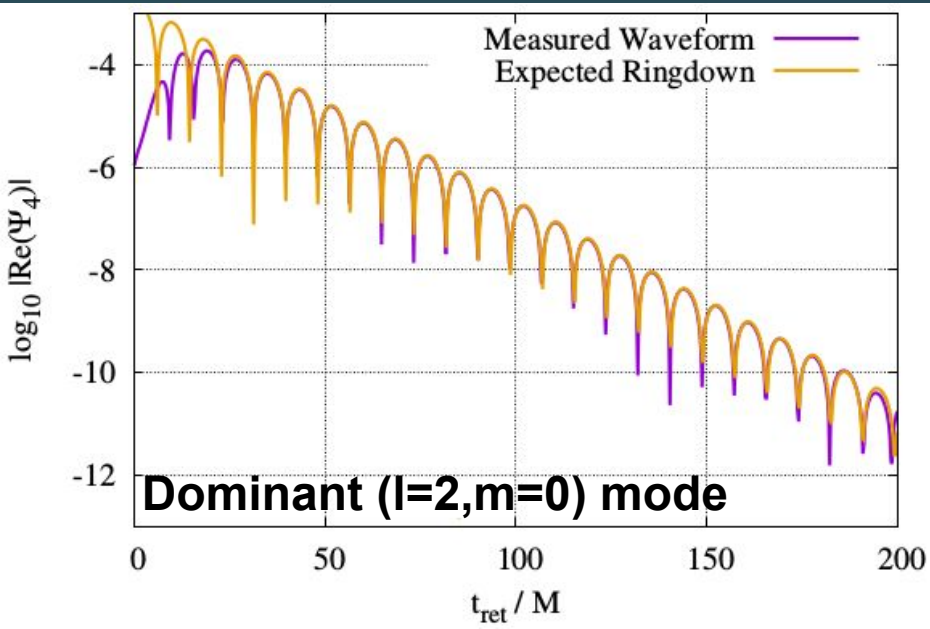
- Dual black hole simulation
 - Brill-Lindquist initial data
 - Moving puncture gauge
 - Sinh-spherical coordinates
 - Moderate resolution



- BH perturbation theory prediction
 - Agreement to ~ 7 decades!

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- Increase FD order, grids fixed
 - Nearly exp. convergence in WFs



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Additional Slides

SENR/NRPy+: Nearly Exponential Convergence

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- When truncation error dominates, finite difference error should scale as

$$\left| C_n u^{(n+1)}(\xi) \right| (\Delta x)^n$$

- If grid is uniform,

$$C_n \sim 1/4^n$$

- Thus *so long as* $u^{(n+1)}(\xi)$ *is bounded*, exponential convergence with *fixed* Δx but increasing FD order n can be expected.
- We find that although this function is not bounded in our cases, we observe *nearly* exponential convergence at the FD orders chosen